Modelling fission gas behaviour in fast reactor (U,Pu)O₂ fuel with BISON

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10 Abstract

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The physics-based fission gas behaviour model available in the BISON fuel performance 11 code provides satisfactory predictive capabilities for application to light-water reactor con-12 ditions. In this work, we present a model extension for application to fast reactor $(U, Pu)O_2$ 13 fuel. In particular, we detail the introduction of a lower bound to the number density of 14 grain-face bubbles, representing a limit to the coalescence process once extensive bubble in-15 terconnection is achieved. This new feature is tested first against an experimental database 16 for UO₂-LWR, and secondly is validated against integral irradiation experiments for fast re-17 actor (U,Pu)O₂ fuel rods irradiated in the FFTF (Fast Flux Test Facility) and in the JOYO 18 reactors. The comparisons of BISON results with the experimental data are satisfactory 19 and demonstrate an improvement compared to the standard version of the code. 20

²¹ Keywords: nuclear fuel, MOX, fast reactors, finite element method, fission gas behaviour

22 1. Introduction

²³ Uranium-plutonium mixed oxide fuels (MOX) are candidate fuels for several fast reactor ²⁴ concepts [1, 2]. Fast reactor conditions pose $(U,Pu)O_2$ fuels under considerable load in terms ²⁵ of high-temperature, steep-temperature gradients and extended burnup [3, 1]. In view of this ²⁶ difference in the operation conditions, the behaviour of fission gas in fast reactor $(U,Pu)O_2$ ²⁷ fuels is different from that of UO₂ in light-water reactors (LWR). [4, 5, 6].

A sufficiently high-temperature gradient can induce the formation of a central void in the fuel pellet [3, 7]. The central void forms due to the migration up the temperature gradient of porosity left over from sintering [3, 7, 8]. Also, the migration of the pores leads to fuel restructuring, exhibiting the formation of new regions within the fuel. In particular, besides the central void, we can identify three other main regions within fast reactor $(U,Pu)O_2$ fuels (Fig. 1), characterized by different microstructures [7]:

In the inner region of the fuel column where the temperature is high enough, the fuel undergoes restructuring driven by an evaporation/condensation mass transport across fabrication pores [3, 1, 9, 10]. In the restructured area, columnar grains are formed, exhibiting lower grain-boundary retention of fission gas than in the unrestructured region because of the cylindrical shape of the columnar grains.

• In the intermediate region, after the restructuring, the microstructure of the fuel is characterized by equiaxed grains [3, 1].

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• The outer region is unaltered. In this outer region, the pores do not move because of the low temperature.



Figure 1: Diagram of microscale features where the region after restructuring are distinguished. For more detailed images and information, see Ref. [11].

Due to this difference in microstructure and in consideration of the high-temperature 43 level in the fast reactor $(U, Pu)O_2$ fuel, attempting to directly apply fission gas behaviour 44 models tailored to UO_2 in LWRs is usually not straightforward [12]. For example, diffi-45 culty arises from the fact that higher temperature implies (exponentially) faster kinetics of 46 diffusive processes governing fission gas behaviour, resulting in a range of model solutions 47 for fission gas release (FGR) and swelling typically not achieved in LWR conditions. It is 48 worth noticing that conservative approaches assuming high temperatures and traditionally 49 applied as design criteria are not acceptable predictive models to be applied in fuel perfor-50 mance simulations (e.g., used for the conservative design of the plenum free volume). It 51 has been shown with best-estimate calculation (e.g., see Ref. [13]) that operational fission 52 gas release in prototype lead-cooled and lead-bismuth-eutectic-cooled fast reactors is in the 53 range of 20 to 60%, i.e., much higher than typical values for LWRs but well below the 54 80-90% observed in sodium-cooled fast reactors for power production [1]. Also, the range 55 of 80–90% of release refers to past generations of sodium-cooled fast reactors, because of 56 their higher temperature target, while, for the new generation sodium-cooled fast reactors, 57 it is lower. On the other hand, Lead fast reactors (LFRs) and Lead Bismuth eutectic fast 58 reactors (LBE-FRs) operates at lower temperatures; therefore, their FGR is also lower. For 59 these reasons, the development of predictive models for fission gas behaviour in fast reactor 60 $(U, Pu)O_2$ fuel is a key point for the effective application of fuel performance codes, hence, 61 it is targeted by several research programs, like [14, 15, 16]. 62

In this work, we present a version of the model by Pastore et al. [17, 18], which was originally developed and validated for LWR UO₂, adapted to the description of fission gas behaviour in fast reactor (U,Pu)O₂. In particular, we propose the introduction of a

physics-based limit to the process of coalescence of growing grain-boundary bubbles. We 66 implemented the modified version of the model into the BISON fuel performance code [19] 67 and we verified, for local ¹ analyses, the behaviour of this extension to the model against the 68 experimental database presented by Baker [20, 21]. Subsequently, we validated the extended 69 version of the model against integral irradiation experiment performed in fast reactors. It is 70 worth noting that there are other phenomena linked to the FGR that do not depend strictly 71 on the diffusion and the evolution of the gas bubbles, but they are connected to the fuel 72 stoichiometry, the Plutonium content and species mobility. Also, the gas diffusion coefficient 73 used in this work, i.e. Turnbull [22], may not be reliable because it has been developed for 74 UO_2 at LWR conditions. For these reasons, we know that all these phenomena that are 75 peculiar for $(U, Pu)O_2$ irradiated in fast reactors and the model of Pastore [18], tailored on 76 UO_2 at LWR conditions, does not perceive. Therefore, the scope of this work is to show 77 that a model for fission gas behaviour developed for LWR conditions with few improvements 78 could have a good reproducibility of the results for $(U,Pu)O_2$ irradiated in fast reactors. 79 However, this work represents an initial extension and other important phenomena that 80 affect FGR (like stoichiometry, Plutonium content, species migration) will be considered in 81 future work. 82

The outline of this paper is as follows. In Section 2, we briefly summarize the main model equations and we detail the proposed modifications. In Section 3, the modified model is tested for local analyses of UO₂ fuel irradiation at temperatures covering both LWR and FBR (Fast Breeder Reactor) ranges. In Section 4, the modified model is tested for the simulation of integral irradiation experiments under fast reactor conditions with BISON, and results are compared to experimental data.

⁸⁹ 2. Model description

In this section, we briefly summarize the main model equations and we detail the proposed modifications. In the second part, we discuss some fundamental assumptions and parameters that affect the simulated fission gas behaviour.

93 2.1. Model development

According to the model of Pastore et al. [17, 18] fission gas behaviour is described as comprised of two main components, i.e., the intra-granular and inter-granular stages. The intra-granular part of the model can be summarized as follows.

Fission gas atoms generated in the fuel grains diffuse towards the grain boundaries through the processes of trapping into and re-solution from nanometer-size intra-granular gas bubbles. The gas arriving at the grain-faces precipitates into grain-face gas bubbles, responsible for grain-face swelling. Fission gas diffusion from within the fuel grain (assumed as spherical) to the grain boundary in one-dimensional spherical geometry is described by:

$$\frac{\partial C}{\partial t} = D_{\text{eff}} \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial C}{\partial r} \right) + \beta \tag{1}$$

where C (atoms m⁻³) is the intra-granular gas concentration (as single atoms in dynamic solution with the matrix and in bubbles), t (s) is the time, r (m) is the radial coordinate in

¹With local analysis and simulation, we mean a simulation using a simplified fuel-only model with a single-cube mesh, in order to simulate the local behaviour of the model in small fuel samples.

⁹⁶ the spherical geometry, β (atoms m⁻³ s⁻¹) is the gas generation rate, D_{eff} (m² s⁻¹) is the

- ⁹⁷ effective gas diffusion coefficient from Speight [23].
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The inter-granular part of the model involves the concurrent calculation of gaseous fuel swelling due to grain-face bubbles and fission gas release, through a description of grain-face bubble development [17, 18]. In summary, an initial number density of grain-face bubbles, N_0 (bubbles m⁻²), is considered with further nucleation during irradiation being neglected (one-off nucleation). Bubbles are assumed to have lenticular shape of circular projection on the grain faces and to all have the same size at any instant. The absorption rate of gas atoms at bubbles is considered to equal the arrival rate at grain boundaries from Eq. 1. The phenomenon of growth (or shrinkage) of grain-face bubbles is described by Pastore et al. [17] referring to the model of Speight and Beere [24]. The overpressurization of the gas bubbles is the driving force of the vacancy absorption/emission mechanism:

$$\frac{\mathrm{d}n_v}{\mathrm{d}t} = k\left(p - p_{\mathrm{eq}}\right) \tag{2}$$

where n_v (vacancies bubble⁻¹) is the number of vacancies per bubble, k is a kinetic constant depending (exponentially) on temperature and on the geometry of the grain face (see White [25]), p and p_{eq} (Pa) are the pressure of the gas in the bubble and the mechanical equilibrium pressure, respectively. The gas pressure is calculated based on the number of gas atoms and vacancies in each bubble and using the van der Waals equation of state [17]. Bubble growth leads the grain-face bubbles to interconnect and merge into larger but fewer bubbles. This coalescence process is described, using a modified model of White [25], relating the rate of decrease of the bubble number density, N (bubble m⁻²), due to coalescence to the rate of increase of bubble projected area on the grain-face, A (m²)

$$\frac{\mathrm{d}N}{\mathrm{d}t} = \frac{6N^2}{3+4NA} \left(\frac{\mathrm{d}A}{\mathrm{d}t}\right) \tag{3}$$

Hence, the coalescence process results in progressively fewer, larger bubbles. When the bubbles reach a sufficient level of interconnection, the grain face is considered as percolated and the fission gas is released to the fuel rod free volume through free pathways made of interconnected bubbles. In the model of Pastore et al. [17], this is underpinned by a principle of grain-face saturation: after the fractional coverage F = NA has reached the saturation value F_{sat} , further bubble growth is compensated by gas release in order to maintain the constant coverage condition:

$$\frac{\mathrm{d}n_{\mathrm{FGR}}}{\mathrm{d}t} = 0 \quad \text{if} \quad F < F_{\mathrm{sat}}$$
$$\frac{\mathrm{d}n_{\mathrm{FGR}}}{\mathrm{d}t} = n\frac{N}{A}\frac{\mathrm{d}A}{\mathrm{d}t} \quad \text{if} \quad F = F_{\mathrm{sat}} \tag{4}$$

where n_{FGR} (atoms m⁻²) is the number of gas atoms released to the fuel rod free volume per unit grain-boundary surface and n (atoms bubble⁻¹) the number of fission gas atoms per grain-face bubble. The saturation point is assumed as a constant with a value of $F_{\text{sat}} = 0.5$ [25, 17, 18].

¹⁰³ The model by Pastore et al. [17] is currently the default option in BISON [19], and ¹⁰⁴ has been extensively validated in LWR UO₂ cases [17, 18, 26, 27]. It is expected that for ¹⁰⁵ $(U,Pu)O_2$ fuels irradiated in fast reactors this model requires modifications and further ¹⁰⁶ validation. In fact, it has been demonstrated [5, 28] that fission gas release reaches high values (outside of the current validation range of the model) at high temperatures often achieved by $(U,Pu)O_2$ fuels irradiated in fast reactors. In these conditions, the kinetic constant k in Eq. 2 becomes large, resulting in bubbles rapidly growing by diffusion-driven absorption of gas atoms and vacancies. As a consequence of bubble growth, the bubble number density strongly decreases, governed by Eq. 3.

Nevertheless, there should be a physical limit to the coalescence process, i.e., bubble 113 growth would not result in noticeable further reduction of the number of bubbles once ex-114 tensive interlinkage is attained. In a simple approach, this limit can be interpreted as a 115 lower bound to the number density of grain-face bubbles. In the original BISON model, 116 such a lower bound was set to a constant value of 10^{10} bubbles m⁻² [18]. However, this 117 limit is of low importance for LWR fuel calculations, where it is normally not attained. On 118 the contrary, for the analysis of higher temperature fuel such as $(U,Pu)O_2$ fuels irradiated 119 in fast reactors in which bubble growth and coalescence proceed more rapidly, the limit 120 becomes highly important in determining the calculated FGR and gaseous swelling. 121 122

In this work, we further develop the concept of a lower bound for the number density of grain-face bubbles and propose a physically grounded approach to its determination.

Assuming a spherical grain of radius a (m), it has a surface area of $4\pi a^2$ (m²). Idealizing the surface area as divided into a number of grain faces n_{face} , each face has an area of

$$A_{\text{face}} = \frac{4\pi a^2}{n_{\text{face}}} \tag{5}$$

Considering that each grain face is shared by two grains, and considering a number of equivalent bubbles per grain face n_{eq} , we can express the bubble number density as

$$N = \frac{1}{2} \frac{n_{\rm eq}}{A_{\rm face}} = \frac{1}{2} \frac{n_{\rm eq} n_{\rm face}}{4\pi a^2} \tag{6}$$

With this notation, the limit number density of grain-face bubbles per unit surface, N_{lim} (bubbles m⁻²) is

$$N_{\rm lim} = \frac{1}{2} \frac{n_{\rm eq, lim} n_{\rm face}}{4\pi a^2} \tag{7}$$

where $n_{\rm eq,lim}$ (bubbles face⁻¹) is the average number of bubbles on a grain face when the saturation of bubbles is reached. The estimation of this parameter is obtained via a random numerical experiment detailed in the following subsection. As for the number of faces per grain $n_{\rm face}$, it is assumed to be 14, based on the number of faces of a tetradecahedron.

129 2.2. Random numerical experiment for the derivation of the bubble equivalent number

In this section, we describe the process used for the derivation of the bubble equivalent number, $n_{eq,lim}$. In order to determine a range of reasonable values for the bubble equivalent number, we performed a numerical experiment. The goal of this experiment is just the derivation of the bubble equivalent number and not the bubble evolution itself because we are aware of the limitations of this derivation in describing such a complex phenomenon.

¹³⁵ This numerical experiment presents the following characteristics:

• We generate 1000 square images representative of the grain-faces.

• Each grain face is assumed of size $10 \times 10 \ \mu m^2$.

• On each grain face, we generate overlapping single-sized circles of radius R sampled from an uniform distribution $R(\mu m) \sim \mathcal{U}[0, 1]$.

• In each grain face, we count the effective number of bubbles, i.e., the number of not-connected clusters of circles.

For this numerical experiment, we considered that the exact size of the grain face is not relevant for the determination of the proposed lower bound to bubble number density since it is expressed in terms of density². Nevertheless, the assumed size of the grain face corresponds to a grains radius of 10 μ m. This can be determined by approximating the 3D shape of the grain with a sphere of surface $4\pi a^2$. If one assumes it is divided into 14 faces of 100 μ m² each, *a* corresponds to a value of 10 μ m.

Also, both bubbles centres coordinates are sampled from uniform distributions. Bubbles ending up close to the edge are simply trimmed by the edge. We are not interested in catching edge-effects, such as percolation of the grain face, and hence the shape of the edge does not represent a strong limitation to our conclusion.

In principle, given enough trials, the patterns generated with bubbles of different size will be covered also by single-size bubbles (with a higher number of smaller single-size bubbles, i.e., the increase in trials required).

We are aware that overlapping circles is a partial representation of bubble coalescence, missing the relaxation of edges which could be represented by phase-field models [29, 30]. On average, both the phase field and the overlapping circles are a reasonable representation of the experimental data³.

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In Figure 2a, the effective number of bubbles is reported as a function of the fractional 160 coverage F, with each point representing one of the 1000 images generated. The number 161 of effective pores gradually decreases with the increase of the fractional coverage, because 162 of coalescence. As an example, Figure 2b reports one of the images generated, with frac-163 tional coverage F = 0.43 and a number of effective bubbles $n_{eq} = 29$. From Figure 2a, 164 in conditions of saturation, i.e., when $F = F_{\text{sat}} = 0.5$, n_{eq} is in the range 0-20 bubbles 165 face⁻¹. In this work, we assume⁴ $n_{\rm eq,lim} = 10$. Moreover, comparing $N_{\rm lim}$ with the range 166 observed bubble densities at saturation, one can note that the proposed limit in this work 167 is compatible with experimental observations of saturated grain boundaries⁵, see Ref. [25]. 168 169

²This assumption holds if the square is big enough compared to the bubble size.

³This comparison is performed on the vented fraction, and we are careful in claiming its generality, but we believe it is an indication of the viability of our "geometrical" approach.

⁴It should be clarified that the bubble concentration we are considering in our model is independent on the grain size, whereas the absolute number of bubbles is consequently proportional to the grain surface area. This assumption is justified by the fact that we do not consider grain growth and therefore the grain size of the problem is fixed as an initial condition. This modelling choice arises from the fact that classical grain growth models are barely applicable in the temperature range of the fuels considered in this work. This implies that the resulting "grain size" referred to in our work is more a modelling abstract quantity than a physical variable representative of the real grain size.

⁵The values for $N_{\rm lim}$ used in sections 3 and 5 fall in the range of $0.9-270 \times 10^{11}$ bubbles m⁻² observed by White [25]. It is important to keep in mind that the values of $N_{\rm lim}$ (from 5.6×10^{10} bubbles m⁻² to 3.5×10^{11} bubbles m⁻² for a grain radius of 10 and 4 μ m, respectively) are related to saturation conditions, i.e. at a fractional coverage of 0.5 because this model feature becomes relevant when the saturation of grain boundaries has been reached. However, the data by White [25] are taken from fuel irradiated and ramped in conditions which may not be representative of a fast reactor environment, therefore we are inclined to consider this comparison as indicative and not a strong support of the proposed new parameter.



Figure 2: Results of the random numerical experiment used to derive the number of equivalent bubbles per grain face. On the left (a) we report the bubble equivalent number obtained as a function of the fractional coverage, while on the right (b) we report a sample image of a pattern of bubbles covering a grain face as obtained by the random process.

The inclusion in the BISON fission gas behaviour model of a lower bound to the inter-170 granular bubble density bounds the coalescence process and, once the lower limit for the 171 number density of bubbles is achieved, any further bubble growth is compensated by FGR 172 by virtue of Eq. 4, and both bubble number density and size remain constant. Note that in 173 the current model, given the same fractional coverage, the fewer the bubbles the larger the 174 volume they occupy and the higher the total gas atoms they retain [17, 18]. Hence, FGR 175 after the attainment of the limit is expected to be higher compared to the situation where 176 coalescence continues. 177

After the attainment of the lower limit in the number density of grain-face bubbles, 178 in the model, the grain-face bubble population is assumed to reach a stationary situation 179 where further evolution is prevented. Also, after extensive interlinkage, bubbles form elon-180 gated or vermicular structures [25], so that the basic model assumption of circular bubbles 181 becomes stronger in itself. However, the concept of a stationary situation is considered 182 reasonable as an attempt to consider behaviour after extensive bubble interlinkage within 183 the current model framework and in a physically meaningful way. As a final physical justi-184 fication for the introduction in the model of such a lower bound, it should be observed that 185 without any limitation the evolution of grain-face bubbles ignores the fact that bubbles are 186 constrained on a single grain face, and can in principle allow for interconnection between 187 bubbles belonging to different grain faces. 188

189 2.3. Modeling assumptions and parameters

An important parameter that affects FGR is the single gas atom diffusion coefficient. The default intra-granular single gas atom diffusion coefficient used in BISON standard version was provided by Turnbull et al. [31] neglecting the contribution of the athermal term (D_3) . However, for the simulations with the extended model, we decided to use the single gas atom diffusion coefficient proposed by Turnbull et al. [22]. We decided to use the original complete diffusion coefficient proposed by Turnbull [22] as it is since it is the reference coefficient used in most similar models in the state of the art. This diffusion coefficient is described as follows:

$$D_s = D_1 + D_2 + D_3 \tag{8}$$

where D_1 (m² s⁻¹) is the thermal component that depends just on temperature, D_2 (m² s⁻¹) the irradiation-enhanced term that is a function of both temperature and the fission rate and D_3 (m² s⁻¹) is the athermal component depending only on the fission rate. It has been decided to use the coefficient proposed by Turnbull et al. [22] because this diffusion coefficient emphasizes the enhanced and the athermal components dependent on the fission rate.

In addition, another issue related to the simulation of fast reactor $(U,Pu)O_2$ cases with 196 BISON pertains to the grain growth. BISON only incorporates an UO_2 -LWR model [32]. 197 which is not suitable to represent grain growth under FBR conditions. Hence, for this 198 work, because of the lack of a model able to describe the grain growth for these conditions, 199 it has been decided to use constant values for the grain radius⁶. We are aware that the 200 restructuring of oxide fuel in fast reactor conditions determine a dramatic change in the 201 crystalline structure [3, 9], as mentioned in the introduction of this work. Indeed, the 202 inclusion of a consistent model for grain growth, fuel restructuring and their influence of 203 fission gas behaviour will be the subject of future investigations. 204

An investigation about the effects of the grain radius on the limit described in Eq. 7 is performed in Section 5.

207 3. Model test with local simulations with BISON

In this section, we assess the impact of the lower bound of bubble number density 208 introduced via Eq. 7 on the fission gas behaviour model of Pastore et al. [17, 18] at 209 different temperatures. The assessment is carried on analyzing the experimental database 210 from Baker (Ref. [21, 20]). In these experiments, the fuel pins have been irradiated in the 211 UKAEA's Winfrith SGHWR up to a burnup of $\sim 1\%$ and the samples examined by Baker 212 (Ref. [21], [20]) lay in the temperature range from 1273 K up to 2073 K. Therefore, these 213 values represent both operational conditions for UO₂-LWR (up to about 1600 K) and more 214 critical ones (beyond 1700 K), closer to operational conditions of $(U,Pu)O_2$ fuel irradiated 215 in fast reactors. 216

The experimental data from [21, 20] correspond to local information, i.e., small fuel samples from irradiated rods. In order to simulate local behaviour with BISON, we used a single-cube mesh. Every simulation has been performed at a different temperature (starting from 1273 K up to 2073 K, consistent with the Baker database) and assuming a typical value⁷ for the grain radius, i.e., 5 μ m, corresponding to a $N_{\rm lim}$ of 2×10¹¹ bubbles m⁻², coming from Eq. 7.

Figures 3 and 4 report all the results at each of the considered temperatures for the 223 quantities of most interest, namely: the fission gas release FGR (%), the number density 224 of grain boundary bubbles per unit surface N (bubbles m^{-2}) and the volumetric swelling 225 due to grain-boundary bubbles (%). The results are shown for both standard and modified 226 BISON. At lower temperatures (Fig. 3), since the bubble number density at the grain 227 boundaries N does not reach the proposed limit N_{lim} , the resulting FGR and volumetric 228 swelling at the grain boundary are the same in both the cases. In Fig. 4, as temperature 229 raises (from 1773 K and so on), the lower bound for the bubble number density at the 230

 $^{^{6}}$ We assumed the grain radius as constant in time and uniform along the pellet radius. This abstract modelling assumption is to be intended just for the modelling of fission gas behaviour and should not be intended as a general model representative of the fuel microstructure.

⁷This consideration might be correct for equiaxed grains, but here we are not considering the cylindrical grains.

grain boundary proposed in this work is attained (Figs. 3a, 3d, 3g and 3j) and this implies a higher FGR and a lower volumetric swelling at the grain boundary compared to the standard BISON case.

The swelling prediction we observe at high temperature is higher than that observed by 234 White [25]. This is to be expected, since the irradiation histories which end up with swellings 235 around 7% are the result of 5'500 hours of irradiation at 1800°C, while the irradiations 236 considered by White are characterized by slightly higher temperatures but only for ramps of 237 short duration. The irradiation time at high temperature is the principal reason behind the 238 difference in the two values of swelling. In fact, due to the higher temperatures (comparable 239 to ranges for fast reactor $(U, Pu)O_2$ fuel), the associated pronounced bubble growth and 240 coalescence lead to the attainment of the lower bound for the number density of grain 241 boundary bubbles systematically. 242

These results highlight the impact of the proposed model modification at temperatures comparable to fast reactor fuel conditions.



Figure 3: Local simulations on UO₂ from the Baker database ([21] and [20]) performed with BISON in a range of temperatures from 1273 K up to 1673 K. For every temperature the fission gas release FGR (%), the number density of bubbles at the grain boundary per unit surface N (bubbles m⁻²) and the volumetric swelling at the grain boundary (%) evolution as a function of the time are reported. The plots related to N are expressed with a semi-logarithmic scale. In every figure, the purple lines, corresponding to the standard BISON version, cannot be distinguished from the blue ones (with a N_{lim} of 2×10^{11} bubbles m⁻², coming from Eq. 7), because they are overlapped.



Figure 4: Local simulations on UO₂ from the database [21] and [20] performed with BISON in a range of temperatures from 1773 K up to 2073 K. For every temperature the fission gas release FGR (%), the number density of bubbles at the grain boundary per unit surface N (bubbles m⁻²) and the volumetric swelling at the grain boundary (%) evolution as a function of the time are reported. The plots related to N are expressed with a semi-logarithmic scale in order to highlight the differences between the two cases analyzed. The purple line refers to the standard BISON version; the blue line refers to the case that uses the lower bound for the bubble number density at the grain boundary $N_{\rm lim}$ (Eq. 7), corresponding to a value of 2×10^{11} bubbles m⁻².

Rod	FO-2	MK-I at	MK-I at	MK-II
	L09	$50 \mathrm{MW}_{\mathrm{t}}$	$75 \mathrm{MW}_{\mathrm{t}}$	
Fuel Data				
Fuel material	$(U,Pu)O_2$	$(U,Pu)O_2$	$(U,Pu)O_2$	$(U,Pu)O_2$
PuO_2 content (wt%)	26.0	18	18	30
Pellet outer diameter (mm)	5.59	5.4	5.4	4.63
Pellet inner diameter (mm)	1.397	-	-	-
Fuel stack height (mm)	914	600	600	550
Pellet density $(\% TD)$	91.7	93.5	93.5	93
Fuel Rod				
Cladding material	HT-9	316SS	316SS	316SS
Cladding outer diameter (mm)	6.858	6.3	6.3	5.5
Cladding thickness (mm)	0.533	0.35	0.35	0.35
Gap width (μm)	101.6	100	100	100
Nominal plenum height (mm)	1057	600	600	550
Fill gas composition	He	He	He	He
Fill gas pressure (MPa)	0.1	0.3	0.3	0.3
Linear heat rate $(kW m^{-1})$	38.8^{-8}	21	32	40
Max. Burnup (% FIMA)	~ 6	~ 2.5	~ 5	~ 5

Table 1: Fuel pin specifications and irradiation data.

245 4. Integral fuel rod analyses

The integral irradiations simulated⁹ in the present work are taken from the experiments 246 FO-2 [33, 34], MK-I and MK-II [35]. The FO-2 assembly has been irradiated in the Fast Flux 247 Test Facility (FFTF), a sodium fast reactor, for 312 equivalent full power days (EFPD) [33] 248 between December 22, 1984, and April 25, 1986, to a peak burn-up of $\sim 6\%$ FIMA (Fissions 249 per Initial Metal Atom) and a peak fast fluence of 9.9×10^{22} n cm⁻² (E > 0.1 MeV) [34]. 250 This test uses the allow HT-9 as cladding material. The FO-2 assembly is composed by 251 169 fuel pins of twelve different types. The fuel was manufactured with co-precipitated 252 $(U, Pu)O_2$ powder, ensuring a uniform plutonium distribution throughout the pellets [34]. 253 The annular fuel¹⁰ pin L09 was destructively examined. This pin was punctured and it was 254 cut at three different elevations. 255

²⁵⁶ JOYO was the first sodium-cooled reactor leading the FBR development program in Japan.

²⁵⁷ MK-I core has been irradiated in the Japanese experimental fast reactor JOYO at 50 MW_t ²⁵⁸ and at 75 MW_t from 1977 to 1981 [35], up to a burnup of $\sim 2.5\%$ FIMA and $\sim 5\%$ FIMA, ²⁵⁹ respectively. MK-II core was irradiated in the JOYO reactor from 1983 to 1986, up to a ²⁶⁰ burnup of $\sim 5\%$ FIMA. The MK-I and MK-II cores are composed by a maximum of 80 and ²⁶¹ 67 driver assemblies, respectively, surrounded by stainless steel reflectors.

More details about the power histories and other information for the experiments FO-2 and MK can be found in Teague et al. [34] and Shimada et al. [35], respectively. The specifications for these experiments are summarized in Table 1.

 $^{^{8}}$ The provided power refers to a simplified power history from Teague et al. [34] and it is representative of the average value.

⁹Phenomena like oxygen defects are outside of the scope of this work, although their impact on the overall behaviour of fast reactor fuel and on fission gas behaviour is relevant.

5. Calculation results 265

In this section, we show the results of integral rod simulations and a sensitivity analysis 266 about the impact of the grain size on the results. 267

5.1. Integral rod simulation results 268

In this Section, we present the calculation results of integral rod simulations obtained 269 with BISON. In Figs. 5a-5d two different cases have been examined: 270

• A first one that considers the standard version of BISON; 271

• A second case that takes into account the lower limit for the bubble number density 272 at the grain boundary $N_{\rm lim}$ (Eq. 7) dependent on the grain radius proposed in this work with the single gas atom diffusion coefficient presented by Turnbull et al. [22].

As shown in Figs. 5a-5d, one can observe that introducing $N_{\rm lim}$ according to Eq. 7 the 275 fission gas release prediction is closer to the experimental data. These results indicate that 276 the limitation introduced describes better the physical limit of grain-face bubble intercon-277 nection occurring at typical operational conditions for fast reactor $(U,Pu)O_2$ fuel. However, 278 in some cases, the results are not in a good agreement with the experimental data (e.g Fig. 279 5b at a burnup lower than 1% FIMA). In this case, the onset is not well described by the 280 model. The reason could be because further modelling developments are required to take 281 into account other phenomena (like stoichiometry and Plutonium content and distribution) 282 that affect the FGR. 283

Also, the quality of the results could be affected by the use of some parameters that have 284 been developed for LWR conditions, first of all, the gas diffusion coefficient. The relation-285 ship provided by Turnbull [22] could not be enough accurate to describe this phenomenon 286 under such different conditions (e.g. higher temperatures, different microstructure, different 287 stoichiometry). 288

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In order to discuss the effect of grain-size on the results, we propose a sensitivity analysis 290 with an analytical approach. This analysis was performed concerning the model illustrated 291 by Pastore et al. [17, 18]. 292

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5.2. Sensitivity Analysis 294

In this section, we want to add more value to our work, performing a sensitivity analysis study. However, the sensitivity analysis is not the main goal of this work, and this study will be investigated more deeply in future publications. To see the effects on the final release due to the perturbation, generated adding a quantity δC of gas¹¹, it is necessary to linearize¹² the equations (presented by Pastore et al. [17, 18] and seen in Section 2) that govern the fission gas behaviour. Hence, once linearized, Eq. 1 that describes the intra-granular diffusion becomes:

$$\frac{\partial \delta C}{\partial t} = D_{\text{eff}} \frac{1}{a} \frac{\partial^2 \delta C}{\partial r^2} \Big|_a \tag{9}$$

¹⁰Furthermore, the effect of open porosity, due to an annular pellet or to the formation of a central void, is not taken into account because the fission gas behaviour model implemented in BISON operates by point.



Figure 5: Calculation results for FGR compared with experimental irradiation data. The purple line refers to the standard BISON version; the blue line corresponds to the case that uses the lower limit for the bubble number density at the grain boundary $N_{\rm lim}$ (Eq. 7). Fig. 5a refers to the fuel pin L09 from the FO-2 assembly, assuming a constant grain radius of 10 μ m, value reported by [34]. The consequent $N_{\rm lim}$ corresponds to a value of 5.6×10^{10} bubbles m⁻². Figs. 5b-5c show the results for MK-I core irradiated at 50 MW_t and at 75 MW_t respectively, while the simulations for MK-II core are illustrated in Fig. 5d. Figs. 5b-5d, related to JOYO reactor, refer to a constant grain radius of 8 μ m [36], with $N_{\rm lim}$ corresponding to a value of 8.7×10^{10} bubbles m⁻².

Eq. 9 shows an inverse proportionality between the diffusion phenomenon and the grain radius *a*. In fact, the larger the grain, the longer the average diffusion path to the grain boundary.

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After linearizing all the equations governing the grain boundary gas behaviour (Eqs. 2-4) presented by Pastore et al. [17], referring to the fractional coverage $F = F_{\text{sat}} = 0.5$ (/) as stationary steady-state point and adding a δC as perturbation to the system, one

¹¹The scope of the model is limited to fission gas behaviour and does not consider several fundamental effects yet, which are therefore not included in the sensitivity.

 $^{^{12}}$ The overall model is highly non-linear, and linearization is to be done carefully. In our analysis, the linearization is to be intended in the proximity of a set of the state variables of the model and used only for perturbation purpose. In the proximity of the set of state of variables values, the linearization is thus a legitimate approximation.

obtains:

$$\delta n_{\rm FGR}\Big|_{N_{\rm lim}} = \frac{n_{\rm sp}N_{\rm sp}}{A_{\rm sp}} \left[\left(\omega + \frac{2\pi D_v \delta_g}{\Omega S n_{\rm vsp}} t \right) \left(\frac{\delta C}{3N_{\rm sp}} a \right) \right]^{2/3} \tag{10}$$

where $\delta n_{\rm FGR}$ (atoms m⁻²) is the variation of number of fission gas atoms released to the fuel rod free volume after the perturbation, assuming a constant value of $N_{\rm lim}$, $n_{\rm sp}$ (atoms bubble⁻¹) is the number of fission gas atoms per grain-face bubble in the stationary point, $N_{\rm sp}$ (bubbles m⁻²) is number density of grain-face bubbles per unit surface in the stationary point, $A_{\rm sp}$ (m²) is the projected area of grain-face bubbles, ω (m³) is the van der Waals' volume of a fission gas atom, D_v (m² s⁻¹) is the vacancy diffusion coefficient on grain boundaries, δ_g (m) is the thickness of the diffusion layer in grain boundaries, Ω (m³) is the atomic (vacancy) volume in grain-boundary bubbles, S (-) is a model parameter (that depends on the fractional coverage) calculated in Pastore et al. [25], $n_{\rm vsp}$ (vacancies bubble⁻¹) is the number of vacancies per grain-face bubble in the stationary point, and t (s) is the time. Performing a derivative of Eq. 10 with respect to the grain radius a, one can obtain:

$$\frac{\partial \delta n_{\rm FGR}}{\partial a}\Big|_{N_{\rm lim}} = \frac{n_{\rm sp}N_{\rm sp}}{A_{\rm sp}} \left[\left(\omega + \frac{2\pi D_v \delta_g}{\Omega S n_{\rm v_{sp}}} t \right) \left(\frac{\delta C}{3N_{\rm sp}}\right) \right]^{2/3} \left(\frac{2}{3a^{1/3}}\right) \tag{11}$$

in which it is neglected that the proposed $N_{\rm lim}$ depends on the grain size itself Eq. (7). Eq. 299 10 points out a dependence of the FGR to the rod free volume of $a^{2/3}$. This effect is due to 300 the decrease of the surface-to-volume ratio of the grain with the increase of the grain radius 301 a. Looking at the derivative (Eq. 11) one can observe that this process is limited since the 302 $\partial \delta n_{\rm FGR}/\partial a$ follows a law proportional to $a^{-1/3}$. Hence, with larger grains, the surface-to-303 volume ratio decreases and this increases grain boundary storing capacity. When many gas 304 atoms are stored in the grain boundaries, the release does not depend on the intra-granular 305 diffusion anymore (Eq. 9) but just on the law described in Eq. 10. This can happen just 306 when many gas atoms are stored in the grain boundaries, and only after a certain burnup 307 level. As the release threshold is reached, there is an expected impact of the proposed $N_{\rm lim}$ 308 value on the concentration of gas being released. This impact is quantitatively evaluated 309 by considering the variation of $N_{\rm lim}$ with the grain size a, namely: 310

$$\frac{\partial N_{\rm lim}}{\partial a} = -\frac{n_{\rm eq, lim} n_{\rm face}}{4\pi a^3} \tag{12}$$

After the lower bound to the inter-granular bubble number density is reached, the coalescence process described by Eq. 3 is considered to stop. This calls for a different expression of the linearized equations governing the grain boundary gas behaviour, resulting in

$$\delta n_{\rm FGR} = \frac{n_{\rm sp}}{A_{\rm sp}} \left[\left(\omega + \frac{2\pi D_v \delta_g}{\Omega S n_{\rm v_{sp}}} t \right) \left(\frac{\delta C}{3N_{\rm sp}} a \right) \right]^{2/3} N_{\rm lim}$$
(13)

$$\frac{\partial \delta n_{\rm FGR}}{\partial a} = \frac{n_{\rm sp}}{A_{\rm sp}} \left[\left(\omega + \frac{2\pi D_v \delta_g}{\Omega S n_{\rm v_{sp}}} t \right) \left(\frac{\delta C}{3N_{\rm sp}} \right) \right]^{2/3} \left(\frac{2}{3} a^{2/3} \right) \frac{\partial N_{\rm lim}}{\partial a} \tag{14}$$

Thus after the lower bound $N = N_{\text{lim}}$ is reached, there is an impact of the proposed model parameter on the overall fission gas release. By comparing Eqs. 11 and 14, we can observe that

• The amount of gas released from the grain boundaries is proportional to the proposed lower bound for the bubble density (considering a linear approximation of model equations). Higher values used for this parameter, e.g., considering higher number of bubbles per grain face or higher number of faces per grain, hence imply a higher predicted fission gas release.

• The dependency on the grain size of the proposed lower bound, expressed by Eq. 12, implies that the expected dependency of the fission gas release on the grain radius changes after the lower bound is reached.

Since the grain radius a is the only free parameter involved in the definition of the lower bound N_{lim} , we reported in Table 2 the dependency of the discussed figures of merit on the grain radius itself, clarifying the change in dependency between the predicted fission gas release with/without considering any lower bound (i.e., after/before the lower bound is eventually reached). It is worth noting that after the lower bound is reached, the predicted fission gas release becomes monotonically decreasing with the grain radius since higher grain radius correspond to lower values of the proposed lower bound (Eq. 7).

Table 2: Summary of the sensitivity analysis about the dependency on the grain radius for intra-granular and inter-granular behaviour.

Scenarios	Figure of merit	Dependency on grain radius, a
Intra-granular behaviour	Variation of intra-granular gas concentration, δC	$\delta C \sim a^{-1}$
Inter-granular behaviour w/o proposed limit	Variation of fission gas released, $\delta n_{\rm FGR}$	$\delta n_{\rm FGR} \sim a^{2/3}$
Inter-granular behaviour w/ proposed limit	Variation of fission gas released, $\delta n_{\rm FGR}$	$\delta n_{\rm FGR} \sim a^{-4/3}$

After this analytical study, we also investigated the effect of the grain-size on the results performing additional simulations. Fig. 6 shows a comparison between the case using the current version of BISON and the one with $N_{\rm lim}$ proposed in this work, also considering two different constant values of grain radius: 5 and 10 μ m for the L09 pin; 4 and 8 μ m for MK-I and MK-II cores. These values lie in the range that goes from 4 to 10 μ m referred to a realistic size for the grain characteristic dimension after the restructuring (Ainscough et al. [32]). Looking at Fig. 6, one can make some observations.

• The FGR for the cases using the limitation N_{lim} with a lower grain radius (black line) is higher than in the case with a higher grain radius (blue line). A smaller grain leads to a shorter average diffusion path to the grain boundary.

Hence, due to this intra-granular diffusion effect (Eq. 1), in a smaller grain, more gas atoms will reach the grain boundary and will be released. This effect can be easily observed also in Table 2. In this table, we report the dependency on the grain radius for both the behaviours (intra-granular and inter-granular).

• In Figs. 6a-6b the intra-granular diffusion effect (Eq. 9) prevails again (red line versus purple line) on the surface-to-volume effect described for the grain boundary gas behaviour (Eq. 10), see also Table 2. Hence, in a smaller grain, more gas can reach the grain boundary and then can be released. However, Fig. 6c shows that this is valid only until a certain value of burnup (the purple line crosses the red one). After a certain burnup level, the grain boundary storing capability becomes dominant: many gas atoms are already stored in the grain boundary and ready to be released. As shown by Eq. 10, this storing capability is improved with larger grains.

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• Again up to a certain burnup (approximately ~1.5% FIMA) the case with $N_{\rm lim}$ with a larger grain (blue line) has a lower release compared to the case with no limitations and with a smaller grain (red line). This shows that up to that burnup the intra-granular effect is stronger, but beyond the grain boundary storing capacity becomes the dominant effect. Furthermore, one can notice that the approaching of the bubble number density lower limit, where (from Eq. 7) it corresponds to a value of 2×10^{11} bubbles m⁻², improves the release for the case represented by the blue line. Consequently, once the lower limit for the number density of bubbles is achieved, any further bubble growth is compensated by FGR. This is shown in Fig. 6a and Fig. 6c,



Figure 6: Sensitivity analysis performed varying the grain radius in a realistic range after the restructuring (from 4 to 10 μ m). Fig. 6a refers to the fuel pin L09 from the FO-2 assembly (grain radius of 5 and 10 μ m, with $N_{\rm lim}$ corresponding to 2×10^{11} and 5.5×10^{10} bubbles m⁻², respectively); Figs. 6b-6c show the sensitivity analysis for MK-I core irradiated at 50 MW_t and at 75 MW_t respectively, while the simulations for MK-II core are illustrated in Fig. 6d (grain radius of 5 and 10 μ m, with $N_{\rm lim}$ corresponding to 3.5×10^{11} and 8.7×10^{10} bubbles m⁻², respectively).

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while in Fig. 6b a sufficiently high level of burnup is not reached.

365 6. Conclusion and future work

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The main goal of this work was modelling fission gas behaviour in fast reactor $(U,Pu)O_2$ 366 fuel with BISON. A model for the calculations of fission gas behaviour was already included 367 in BISON. However, this model was validated only for LWR cases so far. We proposed a 368 new physics-based limit to represent conditions of extended bubble growth and interlinkage 369 at grain boundaries that occur at high temperatures (typical conditions achieved during 370 irradiation in FBRs). Therefore, the scope of this work is to show that a model for fission 371 gas behaviour developed for LWR conditions with few improvements could have a good 372 reproducibility of the results for $(U,Pu)O_2$ irradiated in fast reactors. However, we are 373 aware that many other phenomena that affect FGR (like stoichiometry, Plutonium content, 374 species migration) are relevant and they should be considered in future work. We applied 375 the extended model to local simulations (i.e., performed using a simplified fuel-only model 376 with a single-cube mesh) in order to prove the importance of the limit. Hence, we applied 377 the model to integral FBR $(U,Pu)O_2$ fuel rod simulations with BISON demonstrating a 378 marked improvement in the FGR predictions compared to the standard version of BISON. 379 In particular, the results point out that at low temperatures the grain-face bubble growth 380 and the consequent coalescence are far from their physical limits, while at higher ranges 381 those limiting effects occur. Also, the sensitivity analysis performed highlights the strong 382 dependency of the FGR on the grain size. In conclusion, according to the results obtained in 383 this work, the extension to the fission gas behaviour model provides an improved represen-384 tation of fission gas behaviour under fast reactor conditions. Future efforts can be invested 385 in the validation of the proposed model against more irradiation cases, and in the analysis 386 of other kinds of $(U, Pu)O_2$ fuels as well. For example, the model could be improved taking 387 into account the heterogeneity effect of the MOX fuels, due to the presence of plutonium 388 agglomerates, on the fission gas release. This feature becomes very important in the mod-389 elling of MOX fuel irradiated in thermal reactors. This could be done implementing into 390 the present model the outcomes from the studies of Koo et al. [12] and of Ishida and Korei 391 [37]. 392

A further perspective for future work regards the effect of open porosity to the release. The model describes the release by point, but it could be improved to consider also inner surfaces due to the formation of a central void or because the fuel pin is annular.

Another open issue is the development of a grain growth model able to describe better 396 this phenomenon at high temperatures, especially after the restructuring in order to define 397 also the columnar grain region. So far, all the grains have been considered as spheres, but 398 this approximation can be valid only for as-fabricated and equiaxed grain regions. The 399 columnar grains, instead, present a cylindrical shape, hence, in this case, all the equations 400 involved should be described with cylindrical coordinates. This improvement could be also 401 useful for the sensitivity analysis prospect because a grain growth model can feed directly the 402 expression presented in this work for the $N_{\rm lim}$, which can be adopted for a more accurate 403 sensitivity analysis. However, the sensitivity analysis is not the main goal of this work. 404 and this study will be investigated more deeply in future work, taking into account other 405 parameters and phenomena besides the grain size. 406

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