# DESIGN OF OPERATIONAL COMPLIANT TRAJECTORIES THROUGH A HOMOTOPIC DIRECT COLLOCATION ALGORITHM 

Alessandra Mannocchi, Carmine Giordano ${ }^{\dagger}$ and Francesco Topputo ${ }^{\text {* }}$


#### Abstract

Deep-space missions will be performed in the future by several stand-alone CubeSats. For limited budget reasons, these spacecraft need to follow operationalcompliant trajectories: transfers with thrusting and coasting periods imposed at pre-defined time instants. Traditional trajectory optimisation algorithms exhibit convergence problems when handling discontinuous constraints. In this work, a homotopic direct collocation approach is presented: it employs a continuation algorithm that maps the classical bang-bang trajectory of a fuel-optimal low-thrust problem into an operational-compliant solution. M-ARGO CubeSat mission is considered as case study for validation, including a realistic thruster model with variable specific impulse and maximum thrust. The trajectories computed with the developed algorithm are compared with non-operational-compliant solutions. Our algorithm produces transfers similar to the optimal solutions with no operational constraint, both in terms of thrusting profile and propellant mass.


## INTRODUCTION

While traditional missions were designed with large budgets, in the last years we are witnessing a significant reduction of costs. An expression of this trend is the rapid development of interplanetary CubeSat technology. ${ }^{1}$ Several released-in-situ, deep-space CubeSat missions are expected to be launched in the next years by ESA (e.g., LUMIO, ${ }^{2}$ Milani, ${ }^{3}$ Juventas ${ }^{4}$ ). Another class of standalone interplanetary CubeSats will travel to their final destination without the need of a carrying mothership. The Miniaturised Asteroid Remote Geophysical Observer (M-ARGO) ${ }^{5}$ will be the first European CubeSat to perform a similar mission.
The reduced size of miniaturized spacecraft imposes very low budgets in terms of power, propellant, and data handling. ${ }^{6}$ The high specific impulse makes electric propulsion a good candidate for these probes. ${ }^{7}$ Still, electric propulsion requires a significant on-ground flight dynamics effort with regular navigation and guidance operations. To overcome theses issues, spacecraft, and especially CubeSats, will be required to follow operational-compliant (OC) trajectories, consisting of a repetition of a regular pattern of alternating thrusting and coasting arcs (duty cycles). Operations, as communication, ground-based orbit determination and correction, or scientific experiments, will be performed during coasting arcs. They will also ease ground operations, and thus in turn lower the cost of flight dynamics team.
Electric propulsion trajectories are determined through the formulation of a low-thrust optimal trajectory problem (LOTP), a specialisation of the optimal control problem (OCP) for time-continuous

[^0]systems. ${ }^{8,9}$ No analytic solutions exist for this optimal problem, but several numerical techniques, traditionally divided in direct and indirect methods, ${ }^{10}$ have been developed to solve the LOTP. Indirect methods ${ }^{11,12}$ aim at finding the solution of the necessary optimality conditions derived by calculus of variations. Direct methods ${ }^{13,14}$ transcribe the optimal control problem into a parameter optimisation problem, and then use nonlinear programming (NLP) to find the optimal solution. These two methods are usually solved by means of gradient methods, ${ }^{15}$ which require the explicit use of the first and sometimes second order derivatives of the problem. For this reason, the functions of the LOTP are required to be differentiable, and thus, continuous. ${ }^{16}$
Duty cycles are time-dependent and discontinuous constraints, and thus are not straightforwardly introduced in both methods. They yield a small convergence radius, preventing the convergence of the solver to the optimal solution. ${ }^{17}$ Homotopy, or continuation, is a suitable method to deal with discontinuous structures. The homotopic approach allows solving the original, difficult and discontinuous problem, starting from an easier and affordable one. ${ }^{18,19}$ In particular, it has been applied to indirect methods to overcome discontinuity problems as bang-bang control in fuel optimal solutions. ${ }^{20-22}$
In this work, a new technique, called Homotopic Direct Collocation (HDC) algorithm, is derived to generate OC trajectories by enforcing duty cycles through a homotopic approach applied to a Hermite-Simpson direct collocation method, making the problem only gradually discontinuous. This is achieved imposing to a fuel optimal not OC trajectory heavier weights to the intervals corresponding to coasting arcs until an OC trajectory is obtained. In the end, HDC tries to answer to the question whether is possible to map a real-fuel-optimal but not-OC solution, into a fuel sub-optimal but OC solution, and, in case, how to achieve this mapping. To authors' knowledge, the HDC is the first attempt of applying a homotopic approach to a direct collocation algorithm. Usually, the presence of duty cycles is considered in early trajectory design phases with the approximation of imposing a minor value of the maximum thrust available on board. At the contrary, the approach proposed in this work provides directly thrust and the control angles profiles compliant with the duty cycle without any additional hypothesis. Additionally, it is a general approach, able to model any kind of duty cycle that is needed to be imposed.
The results obtained from HDC are applied to the case of M-ARGO mission, as solution to the needs of the trajectory design of this peculiar mission. The interplanetary transfers are computed using a realistic thruster model, which includes variable maximum thrust and specific impulse. HDC solutions are shown to have thrusting profiles, and required propellant mass, similar to the solutions without the duty cycle constraint.

## THE LOW-THRUST OPTIMAL TRAJECTORY PROBLEM

Consider the following dynamics for a spacecraft in spherical coordinates

$$
\dot{\mathbf{x}}(t)=\mathbf{f}(\mathbf{x}, \mathbf{u}, t)=\left[\begin{array}{c}
v_{r}  \tag{1}\\
\frac{v_{\theta}}{r \cos \phi} \\
\frac{v_{\phi}}{r} \\
\mathbf{a}_{S R P}+\mathbf{P} \mathbf{a}_{G}+\mathbf{S}\left[\begin{array}{ll}
v_{r} & v_{\theta} \\
v_{\phi}
\end{array}\right]^{T}+\mathbf{a}_{T} \\
-\frac{T}{I_{s p} g_{0}}
\end{array}\right.
$$

where x is the state vector

$$
\mathbf{x}(t)=\left[\begin{array}{lll}
\boldsymbol{r}, & \boldsymbol{v}, & m
\end{array}\right]=\left[\begin{array}{llllll}
r, & \theta, & \phi, & v_{r}, & v_{\theta}, & v_{\phi}, \tag{2}
\end{array} \quad m\right]
$$

where $\phi$ and $\theta$ are the azimuth and elevation angles in J 2000 reference system respectively, and $m$ is the mass of the spacecraft. The control vector $\mathbf{u}(t)$ is a function of time for low-thrust propulsion models, expressed as

$$
\begin{equation*}
\mathbf{u}(t)=[T, \quad \alpha, \quad \beta] \tag{3}
\end{equation*}
$$

where $T$ is the thrust magnitude, $\alpha$ is the in-plane angle such that $\alpha \in[-180,180]$ deg, and $\beta$ is the out-of-plane thrust angle such that $\beta \in[-90,90]$ deg. The thrust magnitude is equal to the product between the maximum thrust $T_{\max }$ and the throttle factor $u$, such that $u=0$ means null thrust, and $u=1$ means maximum thrust. The gravitational pulling $\mathbf{a}_{G}$ in Eq. (1) describes a full-ephemeris model according to which

$$
\begin{equation*}
\mathbf{a}_{G}=\boldsymbol{g}(\boldsymbol{r})+\mathbf{a}_{G, i}=-\frac{\mu_{\odot}}{r^{3}} \mathbf{r}-\sum_{i \in P} \mu_{i}\left(\frac{\mathbf{r}_{i}}{r_{i}^{3}}-\frac{\mathbf{r}_{i}-\mathbf{r}}{\left\|\mathbf{r}_{i}-\mathbf{r}\right\|^{3}}\right) \tag{4}
\end{equation*}
$$

where $\mathbf{r}$ is the position vector of the spacecraft, $\mu_{\odot}$ is the gravitational constant of the Sun, and thus the first term of Eq. (4) is the primary acceleration due to the Sun. The other terms in the Eq. (4) model the gravitational accelerations given by the 8 planets of the Solar System, represented by the set $P: \mu_{i}$ is the planetary gravitational constant of the $i$-th planet and $\mathbf{r}_{i}$ is the position vector of the spacecraft with respect to it. The numerical formulation of Eq. (4) employs the method by Betts (see Appendix F of Reference 23) to avoid numerical errors in computing the difference between nearly equal numbers. The SRP acceleration $\mathbf{a}_{S R P}$ is expressed as

$$
\begin{equation*}
\mathbf{a}_{S R P}=\frac{Q A}{m} \frac{\mathbf{r}}{r^{3}} \tag{5}
\end{equation*}
$$

where $Q$ is the solar radiation pressure constant and $A$ is the Sun-projected area. The thrusting acceleration $\mathbf{a}_{T}$ is represented by

$$
\mathbf{a}_{T}=\frac{T}{m} \boldsymbol{\alpha}=\frac{T_{\max } u}{m}\left[\begin{array}{c}
\sin \alpha \cos \beta  \tag{6}\\
\cos \alpha \cos \beta \\
\sin \beta
\end{array}\right]
$$

Eventually, $\mathbf{P}$ and $\mathbf{S}$ in Eq. (1) are two rotation matrices defined as

$$
\mathbf{P}=\left[\begin{array}{ccc}
\cos \phi \cos \theta & \cos \phi \sin \theta & \sin \phi  \tag{7}\\
-\sin \theta & \cos \theta & 0 \\
-\sin \phi \cos \theta & -\sin \phi \sin \theta & \cos \phi
\end{array}\right] \quad \mathbf{S}=\left[\begin{array}{ccc}
0 & \dot{\theta} \cos \phi & \dot{\phi} \\
-\dot{\theta} \cos \phi & 0 & \dot{\theta} \sin \phi \\
-\dot{\phi} & -\dot{\theta} \sin \phi & 0
\end{array}\right]
$$

where $\mathbf{P}$ is used to convert $\mathbf{a}_{G}$ in spherical coordinates while $\mathbf{S}$ is used to write the dynamics in a more compact way. The LOTP aims at computing the optimal control function $\mathbf{u}(t)$ that minimizes $J$, the scalar cost function. In space trajectories design, $J$ is commonly represented by the time of flight, ToF, or the propellant mass, $m_{p}$, required to perform the trajectory. When the objective is to save the propellant mass $m_{p}$, the LOTP is called fuel-optimal problem (FOP), and $J$ is referred as $J_{f}$, where

$$
\begin{equation*}
J_{f}=\int_{t_{i}}^{t_{f}} \frac{T(t)}{I_{s p(t)}} \mathrm{d} t \tag{8}
\end{equation*}
$$

There exist several numerical techniques to solve the LOTP, since no analytic solutions exist for it. Direct methods firstly discretised the problem imposing the dynamical constraints and the boundary conditions, and then find the optimal trajectory through the solution of a NLP problem. In particular, direct collocation methods enforce the dynamics with numerical integration schemes, such as Euler or Hermite-Simpson, ${ }^{24}$ allowing the transcription of differential constraints into algebraic ones. In this work it has been employed as starting point the software DIRETTO, ${ }^{25}$ a low-thrust trajectory design tool developed at Politecnico di Milano, which solves the LOTP with a Hermite-Simpson direct collocation method and exploits Ipopt ${ }^{* 26}$ as NLP solver.

## STRUCTURE OF THE OPTIMAL SOLUTION

The necessary conditions for the optimality of the solution of the FOP are derived by introducing the Lagrange multipliers, or costates, $\boldsymbol{\lambda}=\left[\boldsymbol{\lambda}_{r}, \boldsymbol{\lambda}_{v}, \lambda_{m}\right]$ associated to the state $\boldsymbol{x}=\left[\begin{array}{lll}\boldsymbol{r}, & \boldsymbol{v}, & m\end{array}\right]$. Indicating the optimal thrust direction as $\boldsymbol{\alpha}^{*}$, it can be shown ${ }^{27,28}$ that it is

$$
\begin{equation*}
\boldsymbol{\alpha}^{*}=-\frac{\boldsymbol{\lambda}_{v}}{\lambda_{v}} \quad \text { if } \quad \lambda_{v} \neq 0 \tag{9}
\end{equation*}
$$

Inserting this into the necessary conditions for optimality of the LOTP, the Euler-Lagrange equations ${ }^{8,29}$ it can be demonstrated that the optimal throttle factor, $u^{*}$, is determined as

$$
u^{*}= \begin{cases}0 & \text { if } \quad S>0  \tag{10}\\ 1 & \text { if } \quad S<0\end{cases}
$$

where $S$ is the switching function, $S=-\lambda_{v} \frac{I_{s p} g_{0}}{m}-\lambda_{m}+1$.


Figure 1. The typical bang-bang fuel-optimal thrusting profile.
Accordingly to this Pontryagin's maximum principle (PMP), ${ }^{30}$ the fuel-optimal control has a bang-bang profile, having a piece-wise discontinuous structure, either zero or maximum value, as

[^1]in Figure 1. In particular this represents the fuel optimal thrust profile for the transfer of M-ARGO CubeSat to asteroid 2011 MD starting the June 5, 2023, and lasting 830 days (see the following sections for more details on the statement of the M-ARGO scenarios). The solution of a FOP as in Figure 1 is fuel-optimal, but not OC, showing long thrusting and coasting arcs. A deep-space spacecraft can not thrust for long time periods, because it could not perform any other task in the meanwhile, as communicating, controlling its trajectory, and so on. Even for classic spacecraft, thrusting for long time periods means accumulating remarkable errors along the trajectory, and thus, ideally, some coasting phases have to be inserted in between thrusted-arcs.

## THE HOMOTOPIC DIRECT COLLOCATION ALGORITHM

A duty cycle with the duration of $n$ days binds the control in an alternation of thrusting and coasting arcs, with the duration of $n-m$ and $m$ days respectively, as follows

$$
u=\left\{\begin{array}{lll}
{[0,1]} & \text { if } & t \in\left[0, t_{n-m}\right] \quad \text { days }  \tag{11}\\
0 & \text { if } & t \in\left[t_{n-m}, t_{m}\right] \text { days }
\end{array}\right.
$$

where $n$ and $m$ are the design parameters for the duty cycle. Looking at the control profile in Figure 1, it can be noted that it already presents an alternation of thrusting and coasting regimes. Since optimal solutions like this are produced by the solver in order to minimize $J_{f}$, if $J_{f}$ is written such that the control of the $m$-th day of each duty cycle is less convenient in terms of propellant with respect to the controls of the remaining $(n-m)$ days, the optimiser can be driven to exclude it from the optimal thrusting profile. The idea is to overweight in $J_{f}$ the time intervals in which the thrust is not desired by considering penalty factors, or weights. To achieve this, we firstly discretised the integral of $J_{f}$ in Eq. (8) over each time interval of duration $h_{k}$, and then modified it with penalty weights $w_{k}$. For electric thrusters the specific impulse $I_{s p}$ varies with the input power, and so with the distance from Sun, and thus this variation should be considered in the cost function. The controls are linearly interpolated in each segment, such that $T_{k}=\frac{1}{2}\left(T\left(t_{k}\right)+T\left(t_{k+1}\right)\right)$ and $I_{s p, k}=\frac{1}{2}\left(I_{s p}\left(t_{k}\right)+I_{s p}\left(t_{k+1}\right)\right)$.

$$
\begin{align*}
& J_{f}=\sum_{k=1}^{N_{s}} \frac{T_{k} h_{k}}{I_{s p, k}}  \tag{12a}\\
& J_{f}=\sum_{k=1}^{N_{s}} \frac{T_{k} h_{k} w_{k}}{I_{s p, k}} \tag{12b}
\end{align*}
$$

In Eqs. (12) $N_{s}$ is the number of intervals $\left[t_{k}, t_{k+1}\right]$ considered for the Hermite-Simpson collocation. For each of them, weights $w_{k}$ are unitary during thrusting arcs, while they are selected higher than one during coasting arcs.

The discontinuity posed by the instantaneous switching of the engine from on to off and viceversa in a bang-bang control can be difficult to be solved. Consider Figure 2: theoretically, on each node $t_{m}$, corresponding to the end of the non-thrusting regime of a duty cycle, two different values of the control should be considered, one with $u=0$ right before that instant, at $t_{m}^{-}$, and one with $u>0$ right after it, at $t_{m}^{+}$. The same problem, with inverted values of the control, is present at $t_{(n-m)}$, corresponding to the end of the thrusting regime.


Figure 2. Instantaneous switching of the bang-bang control.


Figure 3. Duty cycle with the not-equally distributed time grid.

Since two different values of the control can not be enforced on a single node, but it is allowed to introduce two nodes very close each other, a not-equally distributed time grid has been considered, as in Figure 3. Given the nature of the grid, it is possible to properly model the thrusting and coasting phases. For each duty cycle, the vectors of the time intervals and of the corresponding weights are

$$
\left.\begin{array}{c}
\mathbf{h}=\left[\begin{array}{llllll}
h_{1}, & h_{1}, & h_{1}, & \ldots & h_{1}, & h_{2},
\end{array} h_{3}, \quad h_{2}, \quad h_{3}\right.
\end{array}\right]
$$

The weights have been selected such that $1<w_{a}<w_{c}<w_{b}$. Indeed, since the first interval $h_{3}$ belongs to the last day of thrusting of a duty cycle, it should not be penalised with a weight $w_{a}$ high as the ones in the coasting day.

Similarly, the last interval $h_{3}$ belongs to a coasting day, but since it is linked by the linear interpolation to the control of the first day of the following duty cycle, its weights $w_{c}$ should model the switching-on of the thruster, while not being so penalised as the control weighed by $w_{b}$. Moreover, this selection of the weights ease the convergence, smoothing the discontinuity of the control structure.

Due to numerical noise problem, the weights $w_{k}$ can not be imposed arbitrarily high in order to obtain an OC solution with a single iteration from a non-OC one. This would made the problem almost discontinuous from the very first iteration, so making it difficult to solve. For this reason, starting from a fuel-optimal not-OC solution as initial guess, the weights are introduced gradually


Figure 4. Scheme of the Homotopic Direct Collocation algorithm.
increased at each iteration, up to when an OC solution is obtained. The HDC algorithm explicitly operates as shown in Figure 4. It can be summarised as:

1. firstly, a fuel non-OC optimal trajectory is computed using the Eq. (12a), with an indicative number of nodes, an equally divided grid and a ToF left free to vary, in order to get a good first guess;
2. the above solution is used as the initial guess of a new optimisation using the Eq. (12b), with all weights $w_{k}$ unitary, in order to impose the non-uniform time grid, and with the ToF fixed to the closest integer to the one computed at step 1 . The definitive number of nodes of the grid is selected at this stage accordingly to the value of the ToF in order to have a correct time discretisation. The result is called First Optimal (FO) solution, and it requires the same $m_{p}$ of the first trajectory computed at step 1;
3. the solution of step 2 is selected to be the initial guess of a new optimisation using Eq. (12b), where the weights $w_{k}$ are slightly increased in the time intervals during which the coasting phases are required to be imposed;
4. step 3 is repeated, each time considering the previously computed solution as initial guess and a slightly increased value of the weights in correspondence of intervals representing the coasting phases, until an OC solution is obtained: the final solution is called Homotopic Optimal (HO).

It has been experienced that if too high weights are introduced too early, the optimiser starts having convergence problems with the discontinuity of duty cycles. In case the convergence is not reached, the initial guess is re-fed to the optimiser, employing a slightly smaller value of the weights as compared to the failed value.

## A FIRST EXAMPLE

Considering Figure 1, the long thrusting blocks for the solution of a FOP are yet located in the optimal location given a departure date and a ToF. The HDC idea is that the most convenient duty cycles of a solution both fuel-optimal and OC should be located around them. Thus, an optimal OC
trajectory should be in aspect very similar to the not-OC solution, but with the long thrusting blocks interrupted when the coasting phases are imposed. The authority lost during them is expected to be recovered by HDC with thrusting arcs located immediately before and/or after the long thrusting blocks. To test this intuition, as well as the use of the not-equally distributed grid, a first simple example has been assessed.

| Iteration | $\boldsymbol{w}_{\boldsymbol{l}}$ | $\boldsymbol{w}_{\boldsymbol{h}}$ | $\mathrm{m}_{\mathrm{p}}(\mathrm{kg})$ | $\boldsymbol{\Delta} \mathrm{m}_{\mathrm{p}}(\%)$ |
| :--- | :---: | :---: | :---: | :---: |
| $\mathbf{1 - F O}$ | 1 | 1 | 1.26867 | - |
| $\mathbf{2}$ | 1.15 | 1.35 | 1.28696 | 1.44 |

Table 1. Iterations for the enforced coasting arc of 20 days.

Considering again the interplanetary trajectory of M-ARGO CubeSat to asteroid 2011 MD in Figure 1, we want to force the solver to switch off the engine for 20 days around day 200. In order to achieve this, some weights $w_{l}$, for the day 1 and for the day 20 of the coasting arc, and $w_{h}$, for the remaining 18 days, have been selected such that $w_{l}<w_{h}$ after the first step. Starting from the FO solution in Figure 1, 2 homotopic iterations summarised in Table 1 have been computed to create the 20-days-long coasting arc in the thrust profile.


Figure 5. HDC recover the lost thrust lost changing the control profile.
As it can be noted, the increase in required propellant mass is limited to $1.44 \%$. In Figure 5, the new thrusting profile, in red, has been reported in comparison with the initial FO solution, in black. The coasting arc around the day 200 of transfer is correctly imposed. The solver, to recover the lost thrust, changes the solution globally, elongating the duration of the second thrusting block, and slightly changing the profile around day 400 and day 600 .

## APPLICATION TO M-ARGO MISSION

M-ARGO will be the first stand-alone CubeSat mission designed by the ESA aimed at targeting a near-Earth asteroid. It has the necessity of exploiting OC trajectories, and thus it has been used as study case to test the algorithm. All the assumptions about the thruster and the mission analysis are taken from Reference 5. The CubeSat will have a maximum wet mass of 28.2 kg , with the propellant mass constrained to be lower than 3.4 kg . The Sun-projected area and the reflectivity parameter have been assumed to be $0.3 \mathrm{~m}^{2}$ and 1.3 , respectively. To model the electric thruster it has been assumed that both the maximum thrust, $T_{\text {max }}$, and the specific impulse, $I_{s p}$, depend on the engine input power, $P_{i n}$, which in turn depends on the distance of the spacecraft from the Sun, $r$. This dependence has been modeled with fourth-order polynomials

$$
\begin{gather*}
T_{\text {max }}=a_{0}+a_{1} P_{i n}+a_{2} P_{i n}^{2}+a_{3} P_{i n}^{3}+a_{4} P_{i n}^{4}  \tag{15}\\
I_{s p}=b_{0}+b_{1} P_{i n}+b_{2} P_{i n}^{2}+b_{3} P_{i n}^{3}+b_{4} P_{i n}^{4} \tag{16}
\end{gather*}
$$

$$
\begin{equation*}
P_{i n}=c_{0}+c_{1} r+c_{2} r^{2}+c_{3} r^{3}+c_{4} r^{4} \tag{17}
\end{equation*}
$$

The values of the 15 coefficients $a_{i}, b_{i}$ and $c_{i}$ in Eqs. (15-17) can be found in Reference 5. At 1 AU the maximum thrust is of 1.56 mN , while the specific impulse is of 3582.82 s . To represent the technological limits of the thruster, $P_{i n}$ has been bounded within a minimum, $P_{i n, \min }$, and a maximum, $P_{i n, \max }$, value, 80 W and 130 W respectively.

Initial and final conditions of the LOTP has been retrieved using SPICE Toolkit ${ }^{31,32}$ and the JPL Horizons On-Line Ephemeris System ${ }^{33-35}$ kernels*. The departure is set from the SEL2 point, while the final conditions are set to rendez-vous one of the asteroids selected as possible target of the mission. The initial value for the mass is $m_{i}$, the wet mass of the CubeSat, while its final value $m\left(t_{f}\right)$ is left free to vary to compute the fuel-optimal solution.

$$
\begin{gather*}
\mathbf{r}\left(t_{i}\right)=\mathbf{r}_{L 2}\left(t_{i}\right) ; \quad \mathbf{v}\left(t_{i}\right)=\mathbf{v}_{L 2}\left(t_{i}\right) \quad \text { and } \quad m\left(t_{i}\right)=m_{i}  \tag{18}\\
\mathbf{r}\left(t_{f}\right)=\mathbf{r}_{\text {Ast }}\left(t_{f}\right) \quad \text { and } \quad \mathbf{v}\left(t_{f}\right)=\mathbf{v}_{\text {Ast }}\left(t_{f}\right) \tag{19}
\end{gather*}
$$

As path constraints, it is imposed that the thrust can not exceed its maximum value, that the propellant mass must be positive, and that the ToF can not exceed the maximum value of 3 years, accordingly to the assumption of the mission analysis.

$$
\begin{equation*}
T-T_{\max }(r) \leq 0 ; \quad\left(t_{f}-t_{i}\right)-1095 \text { days } \leq 0 \quad \text { and } \quad-m\left(t_{f}\right) \leq 0 \tag{20}
\end{equation*}
$$

The duration of the duty cycles is usually thought to preserve a standard working week. In order to keep the mission lifetime within a reasonable time span, it is preferable to not allocate less than six days for thrusting, while one day is at least required to perform all the no-thrusting operations. Thus, in this work a duty cycle of $n=7$ days has been chosen, with $m=1$ day for no-thrusting operations, and $(n-m)=6$ days employable for thrusting. For the not equally distributed time grid, with reference to Figure 3, the choice has been $h_{1}=86400 \mathrm{~s}=24 \mathrm{~h}, h_{2}=72000 \mathrm{~s}=20 \mathrm{~h}$ and $h_{3}=14400 \mathrm{~s}=4 \mathrm{~h}$.

| Asteroid | DD | ToF (d) | Solution | $\mathbf{m}_{\mathbf{p}}(\mathrm{kg})$ |
| :--- | :---: | :---: | :---: | :---: |
| 2014 YD | 26 Jul 2024 | 645 | FO | 1.11559 |
|  |  |  | HO | 1.16563 |
| 2011 MD | 05 Jun 2023 | 830 | FO | 1.26867 |
|  |  |  | HO | 1.30534 |

Table 2. Summary of the solutions computed with HDC.

With these constraints, HDC has been used to compute several solutions towards the candidate asteroids. In this work, two of them, towards asteroids 2011 MD and 2014 YD, are presented. They are summarized in Table 2, where they are reported the Departure Date (DD), the ToF, and the value of the necessary propellant mass $m_{p}$ for each FO and HO solutions.

[^2]| Iteration | $\boldsymbol{w}_{\boldsymbol{a}}$ | $\boldsymbol{w}_{\boldsymbol{b}}$ | $\boldsymbol{w}_{\boldsymbol{c}}$ | $\mathrm{m}_{\mathrm{p}}(\mathrm{kg})$ | $\boldsymbol{\Delta} \mathrm{m}_{\mathrm{p}}(\%)$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| $\mathbf{1}-\mathbf{F O}$ | 1 | 1 | 1 | 1.11559 | - |
| $\mathbf{2}$ | 1.02 | 1.1 | 1.07 | 1.11666 | 0.10 |
| $\mathbf{3}$ | 1.05 | 1.2 | 1.1 | 1.12154 | 0.53 |
| $\mathbf{4}$ | 1.25 | 1.4 | 1.3 | 1.13162 | 1.44 |
| $\mathbf{5}$ | 1.35 | 1.5 | 1.4 | 1.13879 | 2.08 |
| $\mathbf{6}$ | 1.45 | 1.6 | 1.5 | 1.15296 | 3.35 |
| $\mathbf{7 - H O}$ | 1.65 | 1.8 | 1.7 | 1.16563 | 4.49 |

Table 3. Iterations for the solution to asteroid 2014 YD.

In the case of the trajectory to the asteroid 2014 YD, for the fixed DD of July 26, 2024, a first search for the solution with free final time led to a solution with a ToF of 645 days. The results of the homotopies are reported in Table 3: the values of the weights for the iterations have been selected by trial and error. It has to be noted that as the weights increase going from the FO to the HO solution, the propellant mass increases too, as the solution is becoming progressively sub-optimal. However, the increase in required propellant mass $\Delta m_{p}$ of the HO is limited up to $4.49 \%$.

In Figure 6, the iterations of the homotopy have been reported: the black profile is the FO solution, the red ones represent the thrusting profile at each iteration, and the yellow one represents the maximum available thrust. As expected, the additional duty cycles are located around the main thrusting blocks of the FO. It is possible to note the spreading of the thrusting blocks as the weights increase, but also the rising of two new blocks at the beginning of the transfer. The first of them is slightly observable in the FO solution, and as the homotopy proceeds it is increasingly exploited. The second one, instead, arises from scratch. It can be also noted that the thrusting block from day 300 to around 450 is the most convenient one, being the last one to be transformed into an OC one.

| Iteration | $\boldsymbol{w}_{\boldsymbol{a}}$ | $\boldsymbol{w}_{\boldsymbol{b}}$ | $\boldsymbol{w}_{\boldsymbol{c}}$ | $\mathbf{m}_{\mathbf{p}}(\mathrm{kg})$ | $\boldsymbol{\Delta} \mathbf{m}_{\mathbf{p}}(\%)$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| $\mathbf{1}-\mathbf{F O}$ | 1 | 1 | 1 | 1.26867 | - |
| $\mathbf{2}$ | 1.02 | 1.1 | 1.07 | 1.27052 | 0.15 |
| $\mathbf{3}$ | 1.05 | 1.2 | 1.1 | 1.27706 | 0.66 |
| $\mathbf{4}$ | 1.25 | 1.4 | 1.3 | 1.29306 | 1.92 |
| $\mathbf{5}$ | 1.35 | 1.5 | 1.4 | 1.30256 | 2.67 |
| $\mathbf{6}-\mathbf{H O}$ | 1.45 | 1.6 | 1.5 | 1.30534 | 2.89 |

Table 4. Iterations for the solution to asteroid 2011 MD.

For what concerns the solution to the asteroid 2011 MD, for the fixed DD, June 5, 2023, the search for the free-time solution led to a trajectory with a value of the ToF of 830 days. The results of the homotopies are summarised in Table 4. The $\Delta m_{p}$ of the HO is of $2.89 \%$, and the $m_{p}$ required is around 1.31 kg . In Figure 7, the iterations of the homotopy have been reported: the optimal profile is mimicked almost perfectly and no additional thrusting blocks appear. However, a sort of merging of the first two thrusting blocks into a unique operational compliant one can be observed.


Figure 6. Iterations for the solution to asteroid 2014 YD.

## CONCLUSIONS

In this work a new Homotopic Direct Collocation algorithm (HDC), based on a Hermite-Simpson direct collocation and a homotopy approach, has been proposed to overcome the issue of imposing the discontinuous constraint of duty cycles to obtain operational-compliant (OC) trajectories. HDC


Figure 7. Iterations for the solution to asteroid 2011 MD.
has been tested within M-ARGO mission context, and has proved that a mapping of the fuel optimal non-OC solutions into fuel sub-optimal OC ones is possible. HDC trajectories show propellant mass values very close to the ones of the solutions without the OC constraint.

This algorithm opens to the possibility of modeling thrusting and coasting phases into the lowthrust control profile of each considered mission case, at the cost of a minor increase in terms of propellant mass. OC trajectories will have great impact especially on CubeSats mission analysis, having them limited propellant and power budgets. Enabling deep-space CubeSats to follow OC trajectories will also pave the way for major autonomy in space, allowing a great reduction in flight dynamics operations costs.

## ACKNOWLEDGMENT

This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme (grant agreement No. 864697).

## REFERENCES

[1] R. Walker, D. Binns, C. Bramanti, M. Casasco, P. Concari, D. Izzo, D. Feili, P. Fernandez, J. G. Fernandez, P. Hager, D. Koschny, V. Pesquita, N. Wallace, I. Carnelli, M. Khan, M. Scoubeau, and D. Taubert, "Deep-space CubeSats: thinking inside the box," Astronomy \& Geophysics, Vol. 59, No. 5, 2018, pp. 5.24-5.30, 10.1093/astrogeo/aty232.
[2] A. Cervone, F. Topputo, S. Speretta, A. Menicucci, E. Turan, P. Di Lizia, M. Massari, V. Franzese, C. Giordano, G. Merisio, et al., "LUMIO: A CubeSat for observing and characterizing micro-meteoroid impacts on the Lunar far side," Acta Astronautica, Vol. 195, 2022, pp. 309-317.
[3] F. Ferrari, V. Franzese, M. Pugliatti, C. Giordano, and F. Topputo, "Trajectory options for Hera’s Milani CubeSat around (65803) Didymos," The Journal of the Astronautical Sciences, Vol. 68, No. 4, 2021, pp. 973-994, 10.1007/s40295-021-00282-z.
[4] H. R. Goldberg, Ö. Karatekin, B. Ritter, A. Herique, P. Tortora, C. Prioroc, B. G. Gutierrez, P. Martino, and I. Carnelli, "The Juventas CubeSat in support of ESA's Hera mission to the asteroid Didymos," Proceedings of the Small Satellite Conference, Utah State University, Logan, UT, 2019, pp. 1-7.
[5] F. Topputo, Y. Wang, C. Giordano, V. Franzese, H. Goldberg, F. Perez-Lissi, and R. Walker, "Envelop of reachable asteroids by M-ARGO CubeSat," Advances in Space Research, Vol. 67, No. 12, 2021, pp. 4193-4221, 10.1016/j.asr.2021.02.031.
[6] T. J. Martin-Mur, E. D. Gustafson, and B. T. Young, "Interplanetary CubeSat navigational challenges," 25th International Symposium on Space Flight Dynamics (ISSFD), Munich, Germany, 2015.
[7] J. C. Pascoa, O. Teixeira, and G. Filipe, "A review of propulsion systems for CubeSats," ASME International Mechanical Engineering Congress and Exposition, Vol. 1, Pittsburg, PA, 2018, p. V001T03A039, 10.1115/IMECE2018-88174.
[8] A. Bryson and Y.-C. Ho, Applied optimal control. London: Taylor \& Francis, 1975, 10.1201/9781315137667.
[9] D. G. Hull, Optimal control theory for applications. Berlin, Germany: Springer Science \& Business Media, 2013, 10.1007/978-1-4757-4180-3.
[10] J. T. Betts, "Survey of Numerical Methods for Trajectory Optimization," Journal of Guidance, Control, and Dynamics, Vol. 21, No. 2, 1998, pp. 193-207, 10.2514/2.4231.
[11] J. A. Kechichian, "Optimal low-Earth-orbit-geostationary-Earth-orbit intermediate acceleration orbit transfer," Journal of Guidance, Control, and Dynamics, Vol. 20, No. 4, 1997, pp. 803-811, 10.2514/2.4116.
[12] C. L. Ranieri and C. A. Ocampo, "Indirect optimization of three-dimensional finite-burning interplanetary transfers including spiral dynamics," Journal of Guidance, Control, and Dynamics, Vol. 32, No. 2, 2009, pp. 445-455, 10.2514/1.38170.
[13] P. J. Enright and B. A. Conway, "Discrete approximations to optimal trajectories using direct transcription and nonlinear programming," Journal of Guidance, Control, and Dynamics, Vol. 15, No. 4, 1992, pp. 994-1002, 10.2514/3.20934.
[14] C. R. Hargraves and S. W. Paris, "Direct trajectory optimization using nonlinear programming and collocation," Journal of Guidance, Control, and Dynamics, Vol. 10, No. 4, 1987, pp. 338-342, 10.2514/3.20223.
[15] D. Morante, M. Sanjurjo Rivo, and M. Soler, "A survey on low-thrust trajectory optimization approaches," Aerospace, Vol. 8, No. 3, 2021, p. 88.
[16] J. T. Betts, Practical methods for optimal control and estimation using nonlinear programming. Philadelphia, PA: SIAM, 2010, 10.1137/1.9780898718577.
[17] T. Li, Z. Wang, and Y. Zhang, "Double-homotopy technique for fuel optimization of power-limited interplanetary trajectories," Astrophysics and Space Science, Vol. 364, No. 9, 2019, pp. 1-12, 10.1007/s10509-019-3637-6.
[18] F. Jiang, H. Baoyin, and J. Li, "Practical techniques for low-thrust trajectory optimization with homotopic approach," Journal of Guidance, Control, and Dynamics, Vol. 35, No. 1, 2012, pp. 245-258, 10.2514/1.52476.
[19] Y. Wang and F. Topputo, "Indirect Optimization of Fuel-Optimal Many-Revolution Low-Thrust Transfers With Eclipses," IEEE Transactions on Aerospace and Electronic Systems, 2022, pp. 1-13, 10.1109/TAES.2022.3189330.
[20] R. Bertrand and R. Epenoy, "New smoothing techniques for solving bang-bang optimal control prob-lems-numerical results and statistical interpretation," Optimal Control Applications and Methods, Vol. 23, No. 4, 2002, pp. 171-197, 10.1002/oca. 709.
[21] R. Epenoy and R. Bertrand, "Optimal control and smoothing techniques for computing minimum fuel orbital transfers and rendezvous," 18th International Symposium on Space Flight Dynamics (ISSFD), Vol. 548, Munich, Germany, 2004, pp. 131-136.
[22] Y. Wang and F. Topputo, "Indirect optimization of power-limited asteroid rendezvous trajectories," Journal of Guidance, Control, and Dynamics, 2022, 10.2514/1.G006179.
[23] H. Curtis, Orbital mechanics for engineering students. Oxford, UK: Butterworth-Heinemann, 2013, 10.1016/C2011-0-69685-1.
[24] F. Topputo and C. Zhang, "Survey of direct transcription for low-thrust space trajectory optimization with applications," Abstract and Applied Analysis, Vol. 2014, 2014, p. 851720, 10.1155/2014/851720.
[25] F. Topputo, D. Dei Tos, K. Mani, S. Ceccherini, C. Giordano, V. Franzese, and Y. Wang, "Trajectory design in high-fidelity models," 7th International Conference on Astrodynamics Tools and Techniques (ICATT), Oberpfaffenhofen, Germany, 2018, pp. 1-9.
[26] A. Wächter and B. Lorenz T, "On the implementation of an interior-point filter line-search algorithm for large-scale nonlinear programming," Mathematical programming, Vol. 106, No. 1, 2006, pp. 25-57, 10.1007/s10107-004-0559-y.
[27] B. A. Conway, Spacecraft trajectory optimization. Cambridge, UK: Cambridge University Press, 2010, 10.1017/CBO9780511778025.
[28] J. M. Longuski, J. J. Guzmán, and J. E. Prussing, Optimal control with aerospace applications. Berlin, Germany: Springer, 2014, 10.1007/978-1-4614-8945-0.
[29] D. S. Naidu, Optimal control systems. Boca Raton, FL: CRC Press, 2002, 10.1201/9781315214429.
[30] D. F. Lawden, Optimal trajectories for space navigation, Vol. 3. Oxford, UK: Butterworths, 1963.
[31] C. H. Acton, N. Bachman, B. Semenov, and E. Wright, "A look towards the future in the handling of space science mission geometry," Planetary and Space Science, Vol. 150, 2018, pp. 9-12, 10.1016/j.pss.2017.02.013.
[32] C. H. Acton, "Ancillary data services of NASA's navigation and ancillary information facility," Planetary and Space Science, Vol. 44, No. 1, 1996, pp. 65-70, 10.1016/0032-0633(95)00107-7.
[33] J. Giorgini, P. Chodas, and D. Yeomans, "Orbit uncertainty and close-approach analysis capabilities of the horizons on-line ephemeris system," 33rd AAS/DPS meeting, Vol. 33, 2001, pp. 58-13.
[34] J. Giorgini and D. Yeomans, "On-line system provides accurate ephemeris and related data," NASA Tech Briefs, NPO-20416, Vol. 48, 1999.
[35] J. Giorgini, D. Yeomans, A. Chamberlin, P. Chodas, R. Jacobson, M. Keesey, J. Lieske, S. Ostro, E. Standish, and R. Wimberly, "JPL's on-line solar system data service," Bulletin of the American Astronomical Society, Vol. 28, 1996, pp. 25-04.


[^0]:    *PhD Student, DAER, Politecnico di Milano, Piazza Leonardo da Vinci, 32, 20133 Milano.
    ${ }^{\dagger}$ Post-Doc Researcher, DAER, Politecnico di Milano, Piazza Leonardo da Vinci, 32, 20133 Milano.
    ${ }^{*}$ Full Professor, DAER, Politecnico di Milano, Piazza Leonardo da Vinci, 32, 20133 Milano.

[^1]:    *https://coin-or.github.io/Ipopt/index.html (Retrieved on October 8, 2020)

[^2]:    *http://ssd.jpl.nasa.gov/?horizons (Retrieved on May 11, 2020 for kernels)

