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Estimation of impact probability of asteroids and space debris through Monte Carlo Line Sampling and Subset Simulation

KePASSA2017 – ESTEC/ESA

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- Problem of space debris and asteroids
 - Common approaches and alternative proposals
- Advanced Monte Carlo techniques
 - Line Sampling
 - Subset Simulation
- Application of the proposed methods
 - Definition of test cases
 - Results, comparisons and comments
- Conclusions and future developments







- 1. Common approach to the problem of space impacts
- 2. Proposed advanced methods

INTRODUCTION

Introduction



Common approach to the problem of space impacts

- Impacts of asteroids and space debris with planets threaten the safety of human activity in space
 - Planetary protection sets special requirements to avoid the contamination of celestial bodies due to man-made debris in interplanetary missions
 - Time periods under study generally span up to 100 years
- Estimating the impact probability for space objects with celestial bodies requires a large number of long-term orbital propagations with standard Monte Carlo Simulations, resulting in high computational cost
 - The amount of simulations required to estimate the probability within a given confidence level increases as the expected probability decreases

Introduction



Proposed advanced methods

More efficient sampling methods may increase the precision of the probability estimate, or reduce the amount of propagations and the computational cost:

- The Line Sampling method probes the impact region of the uncertainty domain by using lines instead of random points
 - The impact probability is estimated via analytical integration, resulting in a more accurate solution
- The Subset Simulation method computes the impact probability as the product of larger conditional probabilities
 - The method progressively identifies intermediate conditional levels moving towards the impact event, reducing the overall number of samples







- 1. Line Sampling
- 2. Subset simulation

ADVANCED MONTE CARLO METHODS



Line Sampling

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The method follows 4 phases:

1. Determination of the "reference direction"

Starting from a reference solution, a Markov Chain is used to find a direction pointing toward the impact region of the domain

2. Mapping onto the standard normal space

Each sample is mapped from the physical coordinates to normalized ones, in order to associate a normal distribution to each of them

3. Line Sampling

For each sample, a line following the reference direction is probed to identify the limits of the impact region

4. Estimation of impact probability

Probability is analytically estimated as the average of integrals of unit normal distribution obtained along each line







Line Sampling



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Line Sampling







Line Sampling



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Line Sampling







Line Sampling





Line Sampling



Line Sampling

The limit values are used to compute one-dimensional impact probability estimates $\widehat{P}^{k}(I)$ along each line as $\widehat{P}^{k}(I) = \Phi(\overline{c_{2}}^{k}) - \Phi(\overline{c_{1}}^{k})$

where Φ represents the CDF of the unit normal distribution

The total probability and the associated variance are estimated as

$$\widehat{P}(I) = \frac{1}{N_{T}} \sum_{k=1}^{N_{T}} \widehat{P}^{k}(I)$$
$$\widehat{\sigma}^{2}\left(\widehat{P}(I)\right) = \frac{1}{N_{T}(N_{T}-1)} \sum_{k=1}^{N_{T}} \left(\widehat{P}^{k}(I) - \widehat{P}(I)\right)^{2}$$







Subset Simulation



The method uses a series of MC Markov Chains in order to cover the initial uncertainty domain and move toward the impact region:

- 1. Generate *N* samples at conditional level (CL) 0 by standard Monte Carlo
- 2. Propagate each sample and identify its minimum distance from the Earth, and sort them according to their distance from Earth
- 3. Identify samples belonging to the next conditional level
 - Consider fixed values of conditional probability and identify the threshold as the $(1-p_0)N$ th element of the list
- 4. Generate (1- $p_0 N$) new samples belonging to that CL by means of Monte Carlo Markov Chain
- 5. Go back to step 2



























Standard MC (CL0) KePASSA2017 – ESTEC/ESA

Subset Simulation

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Impact region I



Advanced Monte Carlo methods



- Procedure is stopped when the threshold for the current conditional level is lower than the minimum distance set for the impact condition
- Given a sequence of intermediate impact regions I₁ ⊃...⊃ I_n = I, the impact probability is estimated as the product of larger conditional probabilities

$$\hat{P}(I) = \hat{P}(I_1) \prod_{i=1}^{n-1} \hat{P}(I_{i+1}|I_i) = p_0^{n-1} N_n / N$$

- $\hat{P}(I_{i+1}|I_i)$ probability of I_{i+1} conditional to I_i
- p_0 predefined conditional probability (equal for every level)
- n number of conditional levels
- *N* number of samples per conditional level
- N_n number of impacting samples at the last conditional level









- 1. Case 1: definition, results and comparison
- 2. Case 2: definition, results and comparison

APPLICATION OF THE PROPOSED METHODS

Premise

The comparison among the different techniques is performed by analysing the following parameters:

 N_{ς}

- Number of initial random samples
- Total number of propagations
- Impact probability estimate
- Associated standard deviation
- Coefficient of variation
- Figure of Merit

 N_{P} $\hat{P}(I)$ $\hat{\sigma}\left(\hat{P}(I)\right)$ $\delta = \hat{\sigma}/\hat{P}(I)$ $FOM = 1/(\hat{\sigma}^{2} \cdot N_{P})$

Notes:

- The computational time was not considered due to the nature of the numerical integration and of the machine, and the number of simulations was preferred instead as reference parameter
- The FOM parameter puts more emphasis on the accuracy of the results than on the computational burden [1]





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Case 1: 2010RF12 – Definition



- 2010 RF₁₂ is a small Near Earth Asteroid (NEA) with currently the highest probability of hitting the Earth (around 6%)
- 2095 fly-by was chosen as a test case due to the high expected impact probability
- LEO crossing (with LEO under 2000 km from Earth's surface) was considered as main event for the application of the methods

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Case 1: 2010RF12 – Comparison of results

Visualisation of initial uncertainty set and impact region in $(\delta a, \delta l)$ plane





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Case 1: 2010RF12 – Comparison of results

	N _S	N _P	$\widehat{P}(I)$	$\widehat{\sigma}$	δ	FOM	$\widehat{\sigma}_{MC}/\widehat{\sigma}$	FOM /FOM _{MC}
Std MCS	1e4	1e4	7.07e-2	2.56e-3	2.76e1	15.22	1.00e0	1.00e0
LS	1e3	~6e3	7.29e-2	1.13e-3	1.56e-2	122.17	2.26e0	8.03e0
SS	1e3	1.9e3	8.01e-2	6.00e-3	7.50e-2	14.62	4.27e-1	9.61e-1

Both methods correctly identify the impact region

 For high values of the expected probability, the two methods are comparable with the standard Monte Carlo in terms of accuracy



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Both methods correctly identify the impact region

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Case 2: Apophis – Definition



- 2036 close approach may present impacts with Earth depending on the uncertainty given by 2009 observations, and was chosen as a test case as more challenging for the application of the methods
- **GEO crossing** was considered as main event for the application of the methods for computational reasons





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Case 2: Apophis – Comparison of results

Visualisation of initial uncertainty set and impact region in $(\delta a, \delta l)$ plane





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Case 2: Apophis – Comparison of results

	N _S	N _P	$\widehat{P}(I)$	$\widehat{\pmb{\sigma}}$	δ	FOM	$\widehat{\sigma}_{MC}/\widehat{\sigma}$	FOM /FOM _{MC}
Std MCS	1e6	1e6	4.70e-5	6.86e-6	6.85e0	2.13e4	1.00e0	1.00e0
LS	1e5	~1e6	5.32e-5	3.45e-7	6.48e-3	8.46e6	1.99e1	3.97e2
SS	1e3	4.6e3	6.10e-5	1.30e-5	2.10e-1	1.28e6	5.27e-1	6.05e1

Both methods become more convenient as the expected probability decreases, outperforming standard MC either in computational time (SS) or accuracy (LS)

- SS provides an estimate of the probability using a much lower number of propagations, with significant reductions in the computational burden, but also a larger variance
- LS is computationally heavier in its current implementation, but offers the highest accuracy, especially when the estimated probability is low



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Case 2: Apophis – Comparison of results

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- 1. Final considerations
- 2. Future developments

CONCLUSIONS

Conclusions





Final considerations

- Line Sampling
 - Advantages
 - Using the same number of random samples, LS can achieve a lower variance of the solution (higher accuracy) thanks to the use of exact integrals on each line
 - Larger efficiency with respect to standard techniques as the impact probability gets lower
 - Current limitations
 - Current implementation supposes a unique impact region with a regular shape
 - Current implementation uses extra evaluations to probe each line, thus decreasing the efficiency of LS

Conclusions





Final considerations

- Subset Simulation
 - Advantages
 - It drives the generation of samples towards the impact region of the initial uncertainty set
 - Larger efficiency with respect to standard techniques as the impact probability gets lower
 - Reduced computational burden with respect to standard MC and LS
 - Current limitations
 - Correlation among conditional samples: possible bias in the estimates
 - Larger variance with respect to LS method

Conclusions

Future developments

- Line Sampling
 - Analysis of the influence of the main implementation degrees of freedom on the accuracy of the results
 - Obtain an analytical expression for the confidence interval in order to determine a minimum number of lines required for a given confidence level
 - Improvement of the zeros computation
 - Application to **planetary protection** studies as part of a Ph.D. research
- Subset Simulation
 - Analysis of the influence of the main implementation degrees of freedom on the accuracy of the results
 - Optimization of the Monte Carlo Markov Chain for samples generation

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